

A photograph of a twilight sky with wispy, white clouds. The sky transitions from a deep blue at the top to a warm orange and red near the horizon. The silhouettes of trees are visible in the foreground at the bottom of the frame.

Investigations of Atmospheric Aerosol based on the Polarization Measurements of the Twilight Sky

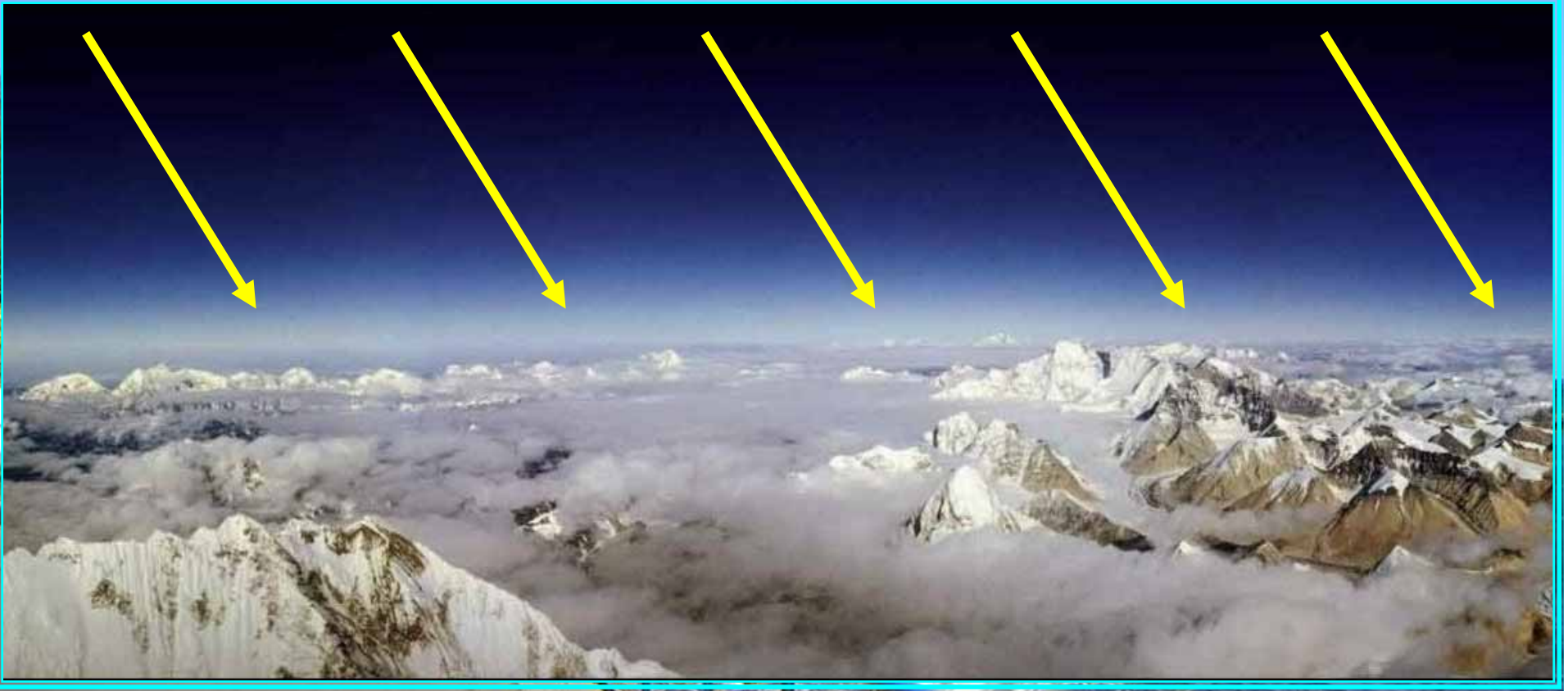
Oleg S. Ugolnikov

*Space Research Institute,
Russian Academy of Sciences*

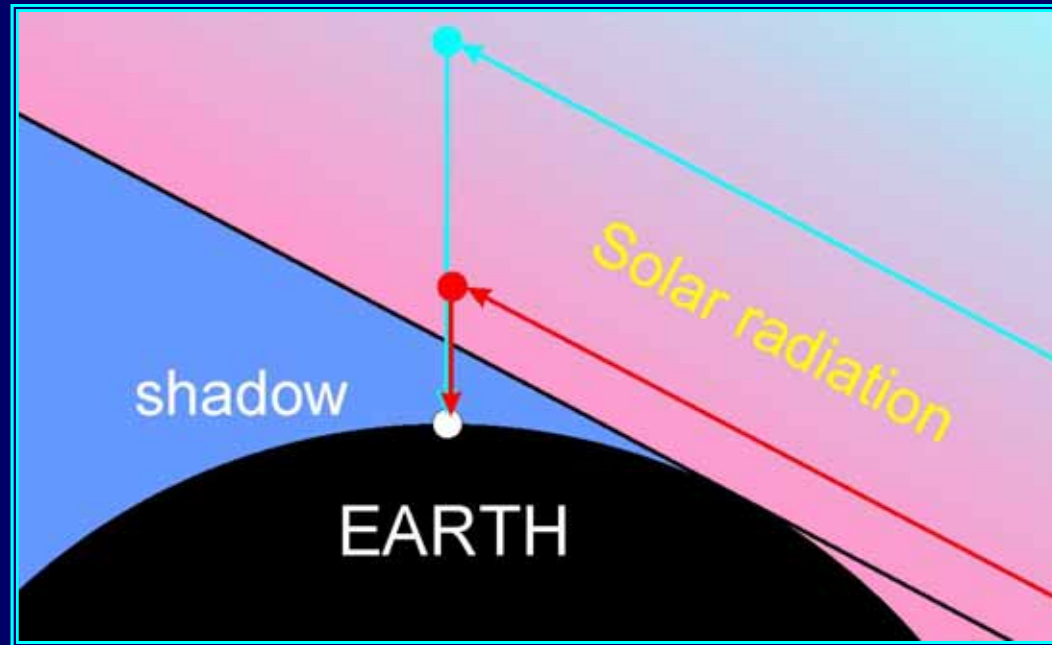
ougolnikov@gmail.com

Daytime sky background

It is defined by the scattering in the near-ground atmosphere (6-8 km)



Altitude scanning



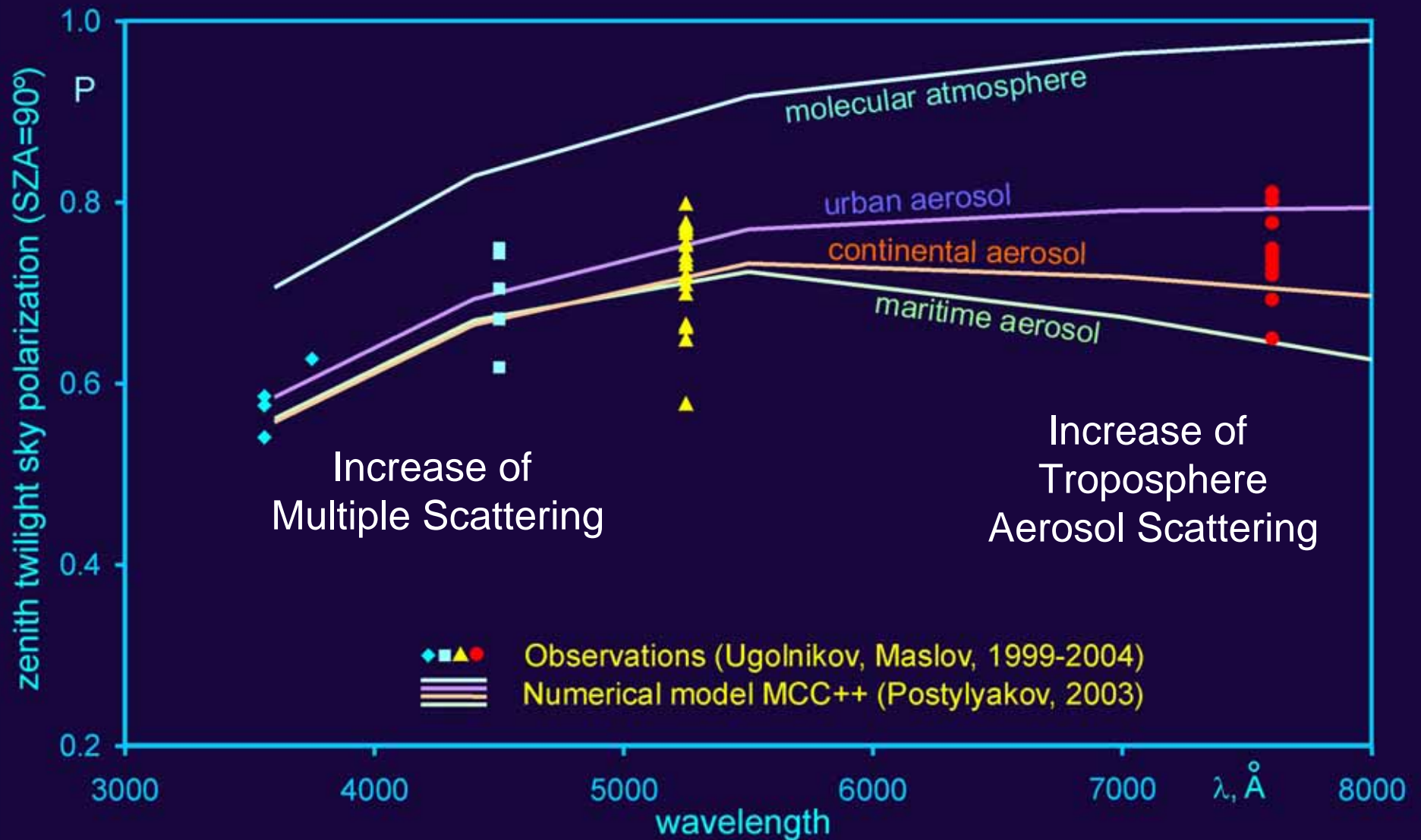
Use of different wavelengths: additional altitude separation, different mechanisms of scattering (molecular, aerosol, resonance, multiple).

Polarization measurements

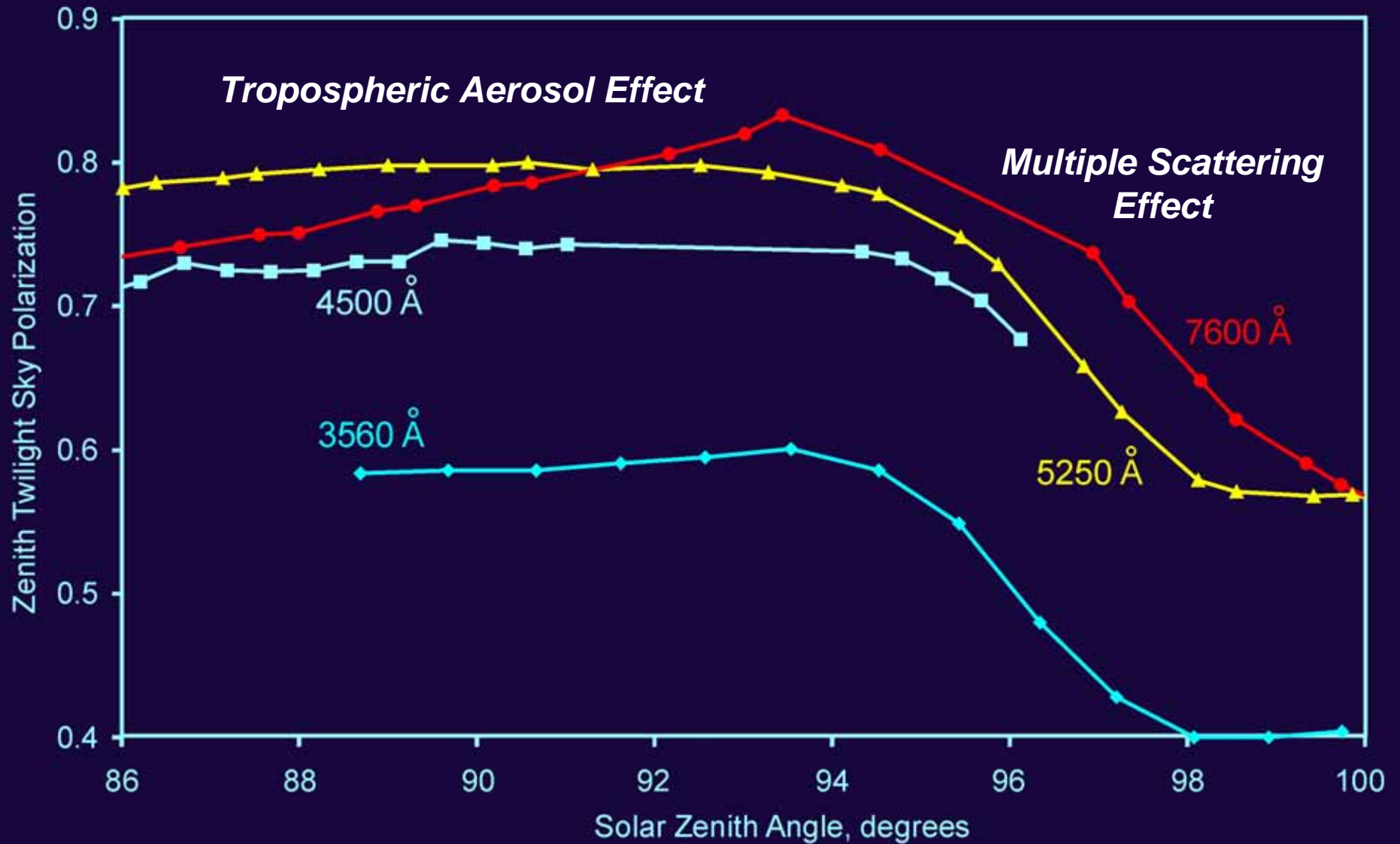
*Help to separate different components of the twilight sky background (molecular scattering, aerosol scattering, **multiple scattering**, ...)*

Expand the information about the atmospheric aerosol (polarization scattering matrix).

Twilight sky polarization (zenith, sunrise)



Twilight sky polarization



Detection of aerosol in the different layers

Observations: Wide-angle CCD-camera with polarization filter (Crimea, Ukraine).

Observable effects: changes of sky brightness and polarization distributions near the zenith for the definite stages of twilight.



Output data:

The value of aerosol scattering contribution in the sky background;

The ratio of aerosol and molecular scattering coefficients;

The polarization of aerosol scattering (scattering angle about 90°).

Stratospheric aerosol

Altitude – above 20 km (the Junge layer).

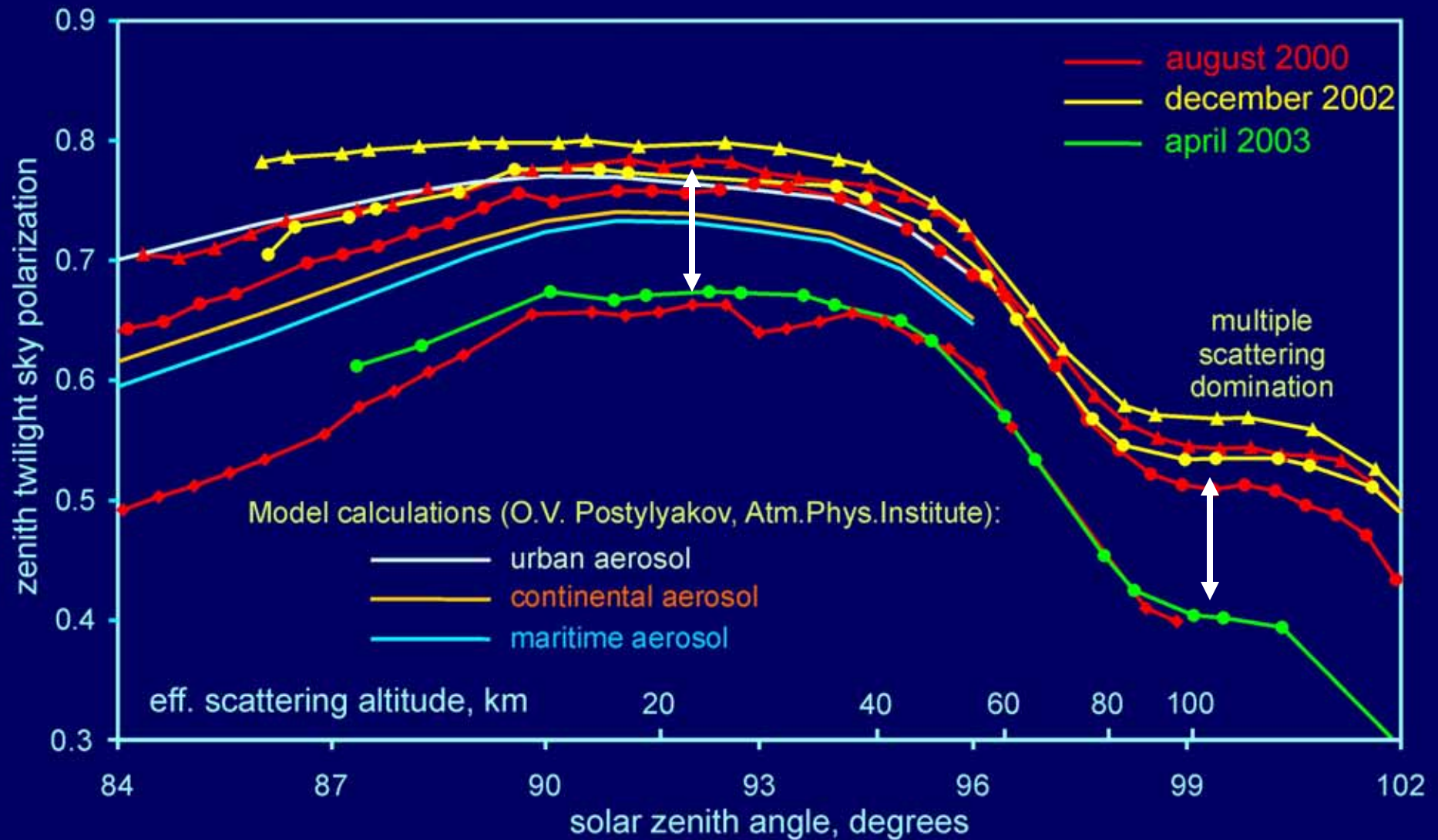
Reveals itself at solar zenith angle 90° - 94° (for 5250 Å).

High latitudes: Polar Stratospheric Clouds.

*Other locations: observed after the strong volcanic eruptions, the strongest one during the last decades was **Pinatubo** (1991, Volcano Eruption Index = 6).*

The stratospheric aerosol level was low since 1995 until the late 2006.

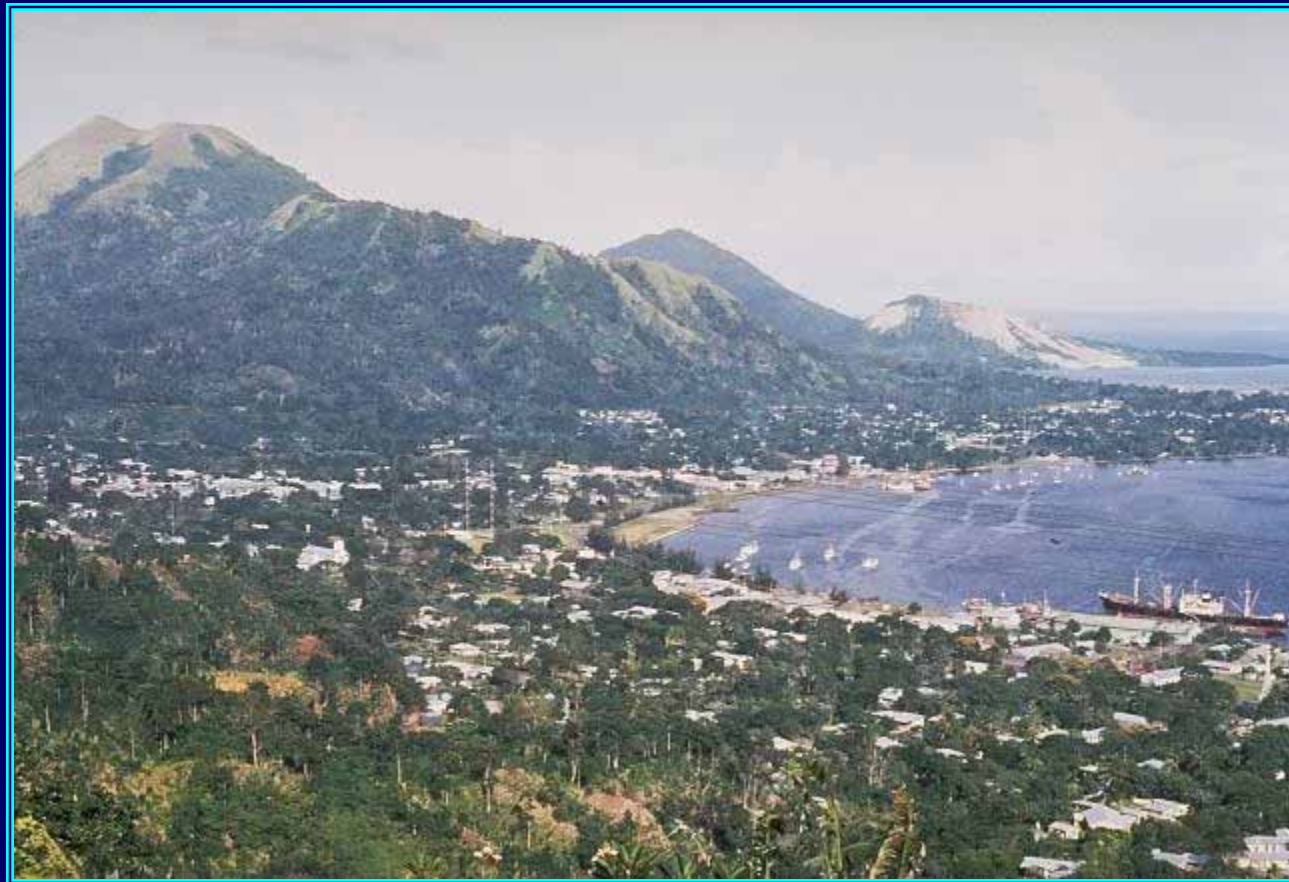
Zenith twilight sky polarization (525 nm) Theory and observations – clear stratosphere



Rabaul Volcano (New Guinea)

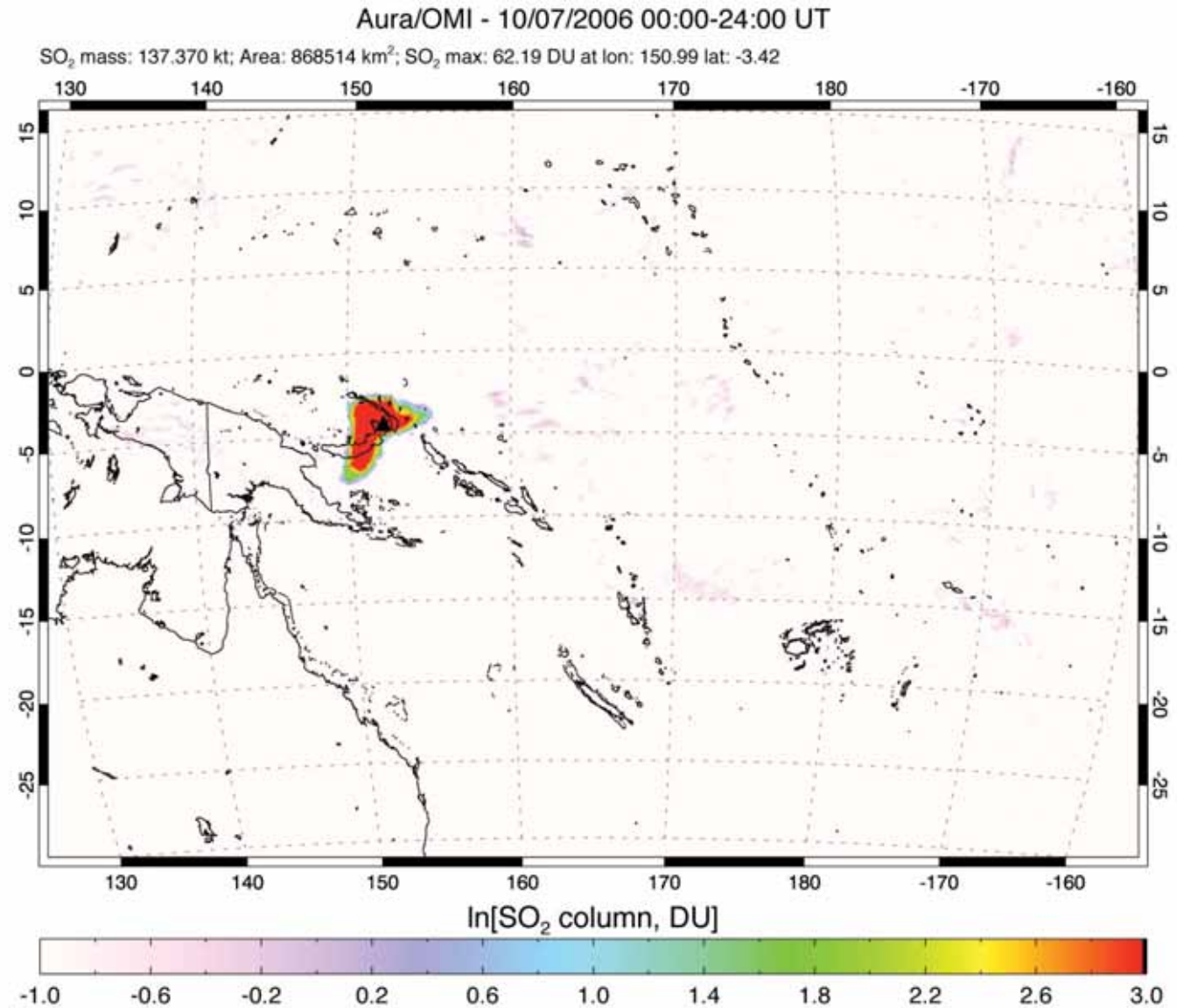
Eruption with index 4 had started at October, 7, 2006.

The eruption products reached the altitude 18 km and threw up to the stratosphere.



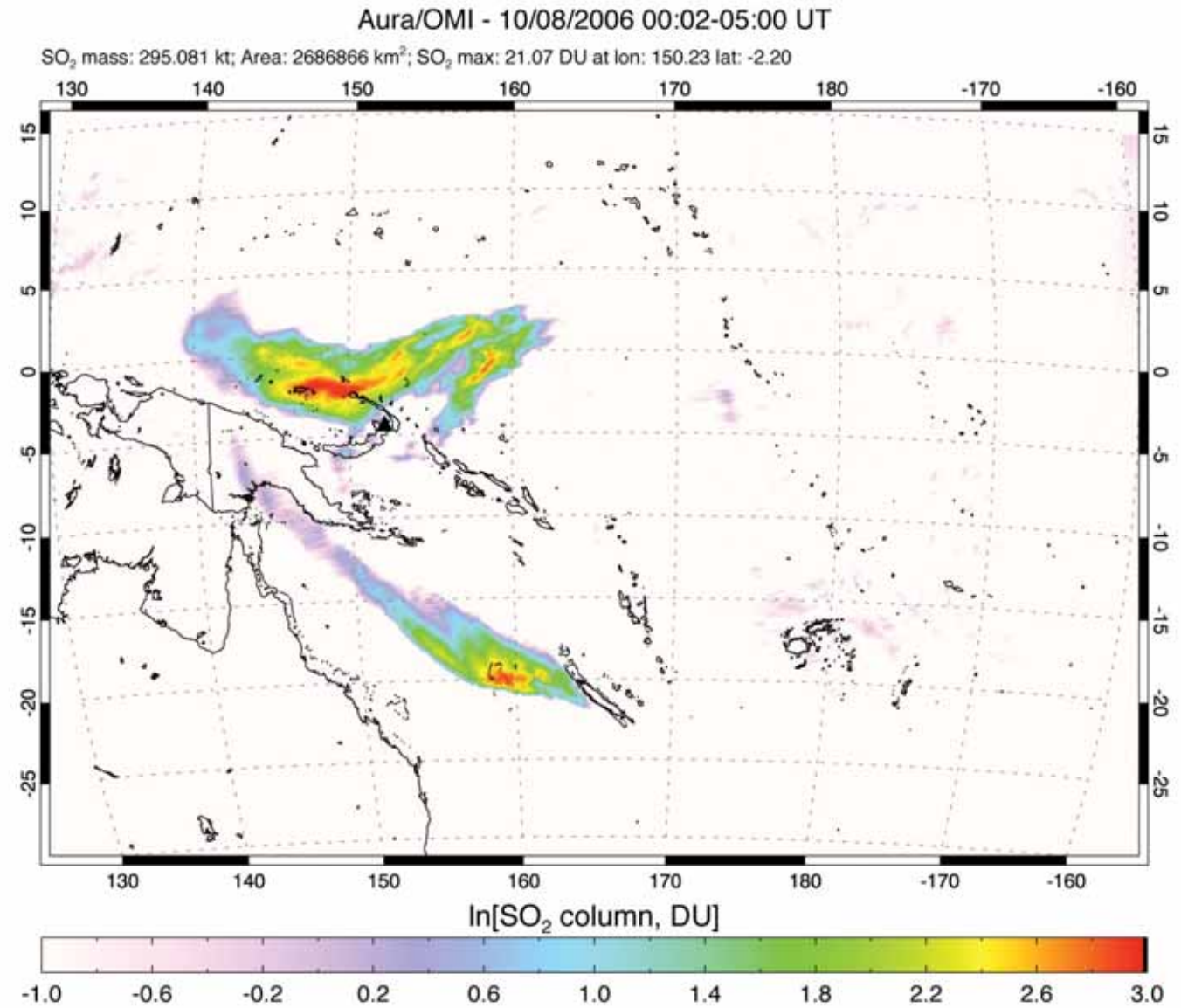
Sulfur dioxide emission

7.X.2006



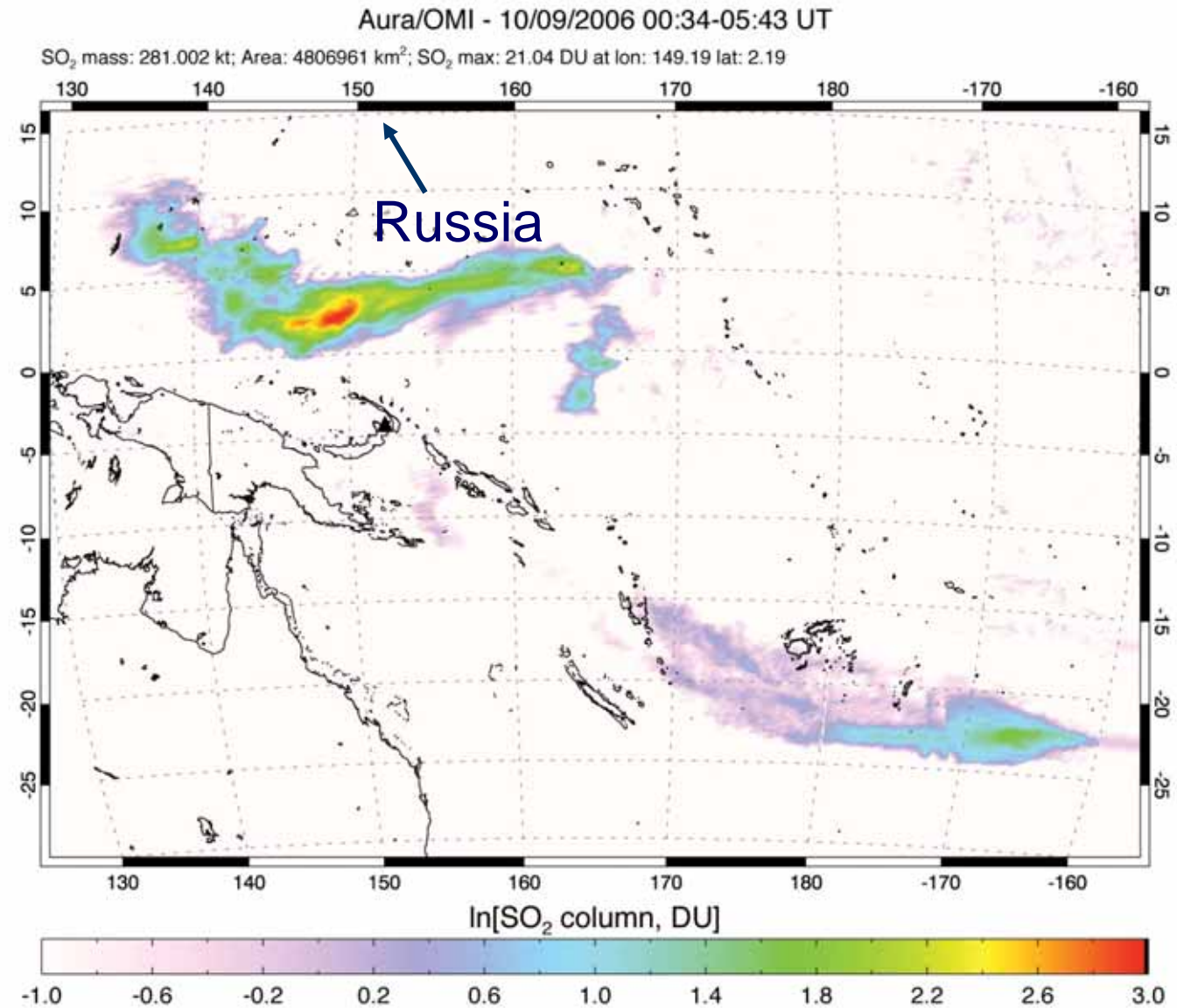
Sulfur dioxide emission

8.X.2006



Sulfur dioxide emission

9.X.2006

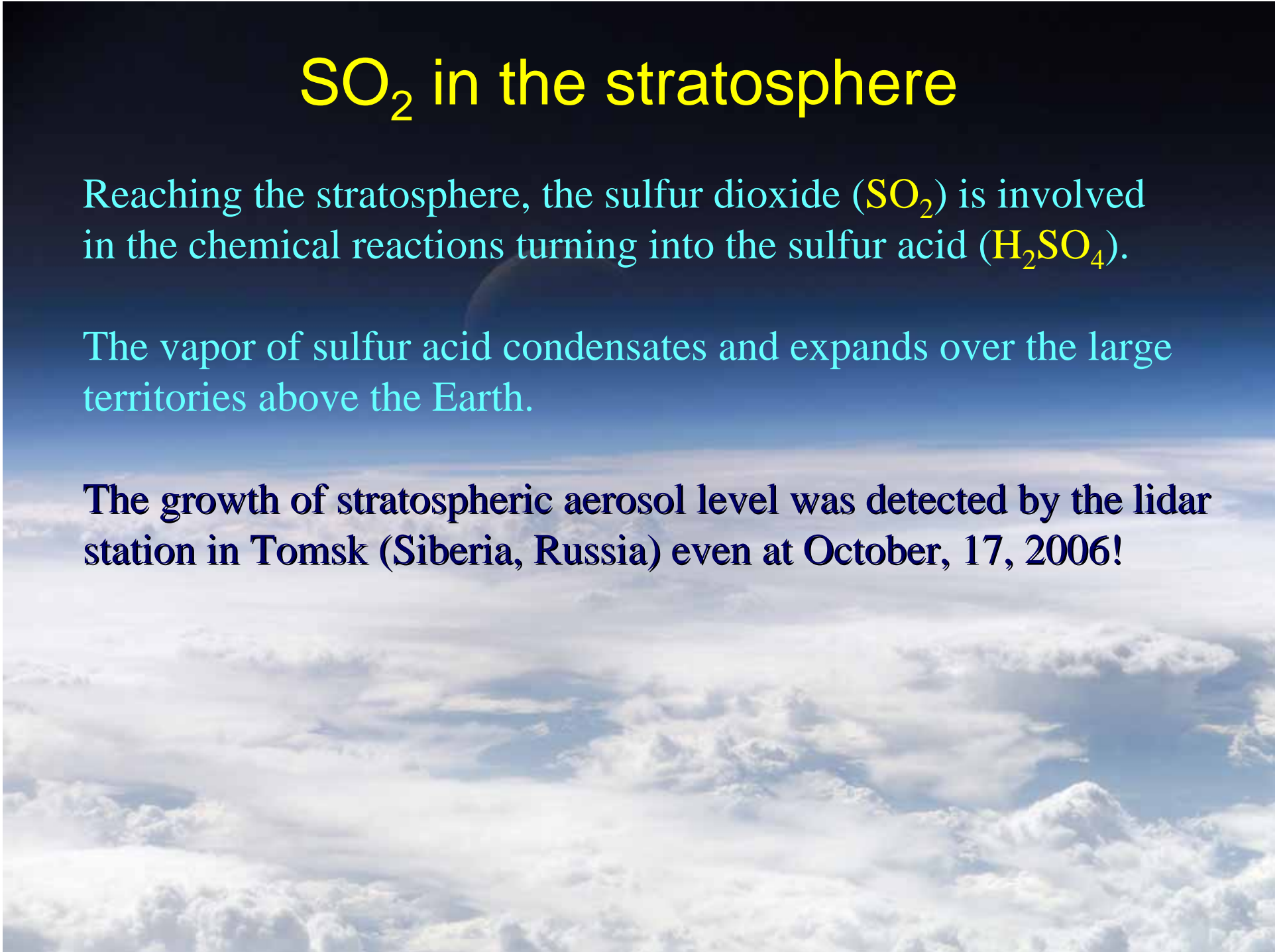


SO₂ in the stratosphere

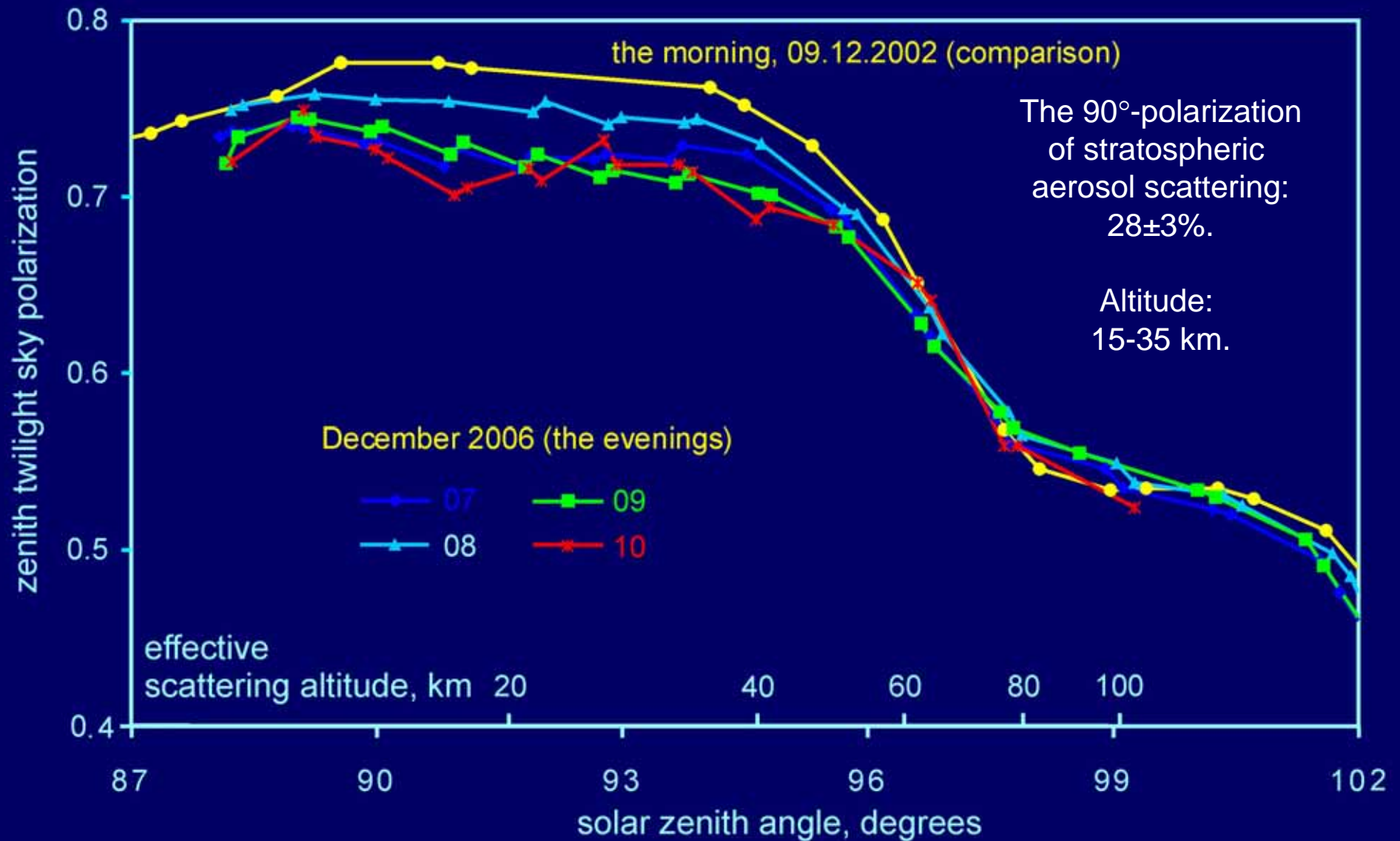
Reaching the stratosphere, the sulfur dioxide (SO₂) is involved in the chemical reactions turning into the sulfur acid (H₂SO₄).

The vapor of sulfur acid condensates and expands over the large territories above the Earth.

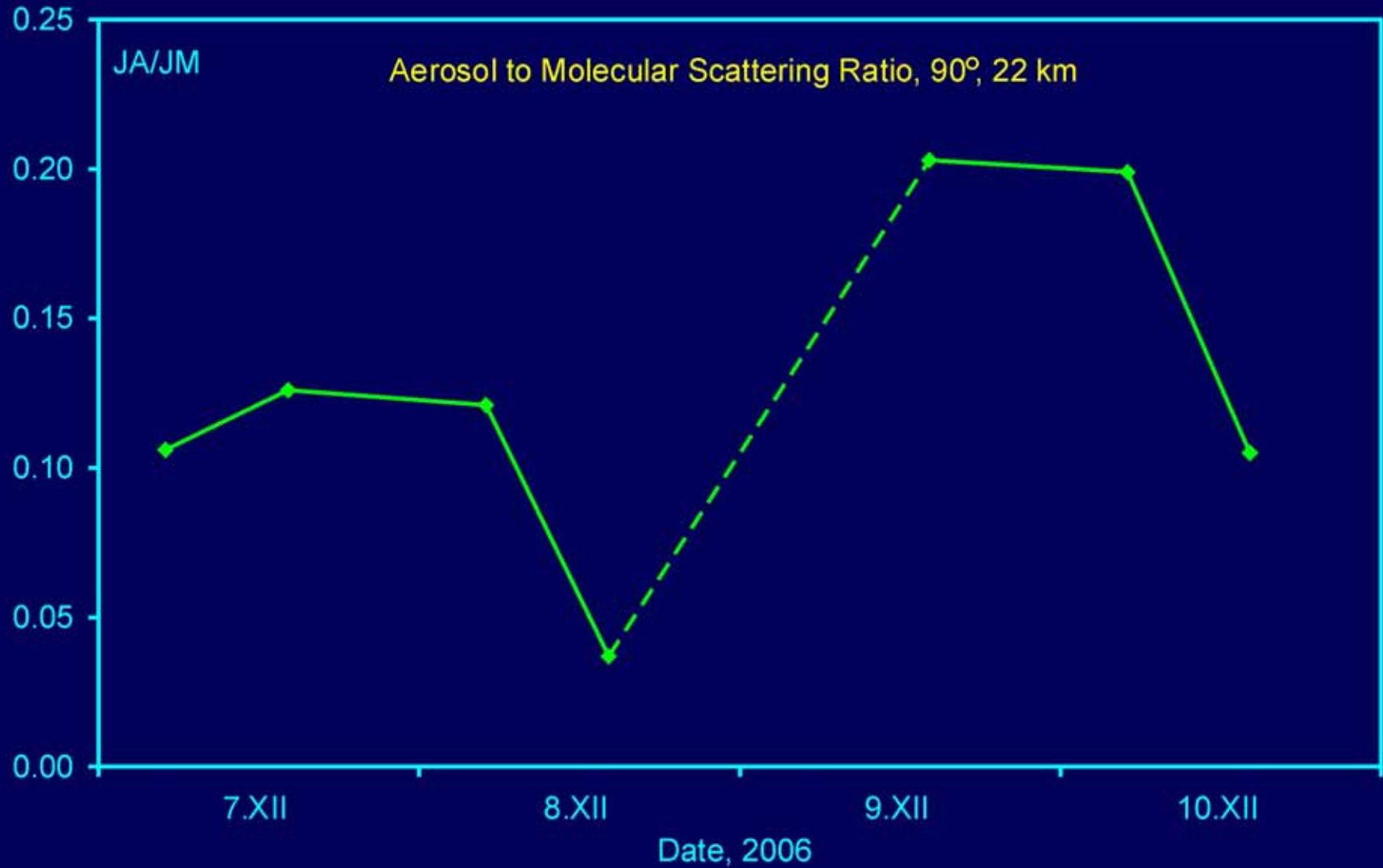
The growth of stratospheric aerosol level was detected by the lidar station in Tomsk (Siberia, Russia) even at October, 17, 2006!



Moonless twilight sky polarization – 525 nm, December, 2006, Crimea



Scattering Coefficients Ratio



Mesospheric Aerosol

*Noctilucent clouds – highest in the atmosphere (80-90 km).
Observed at solar zenith angles more than 96° - 97° .*

*Meteoric dust – appears basically after the maxima of
major meteor showers. It can be the condensation nuclei for
the noctilucent clouds.*

*Twilight photometry had registered the impacts from Quadrantids
(Link, Robley, 1971), Orionids and h-Aquarids (Mateshvili et al.,
1998-1990s).*

Leonids

Parent comet is 56P/
Tempel-Tuttle.

Period: 33 years.

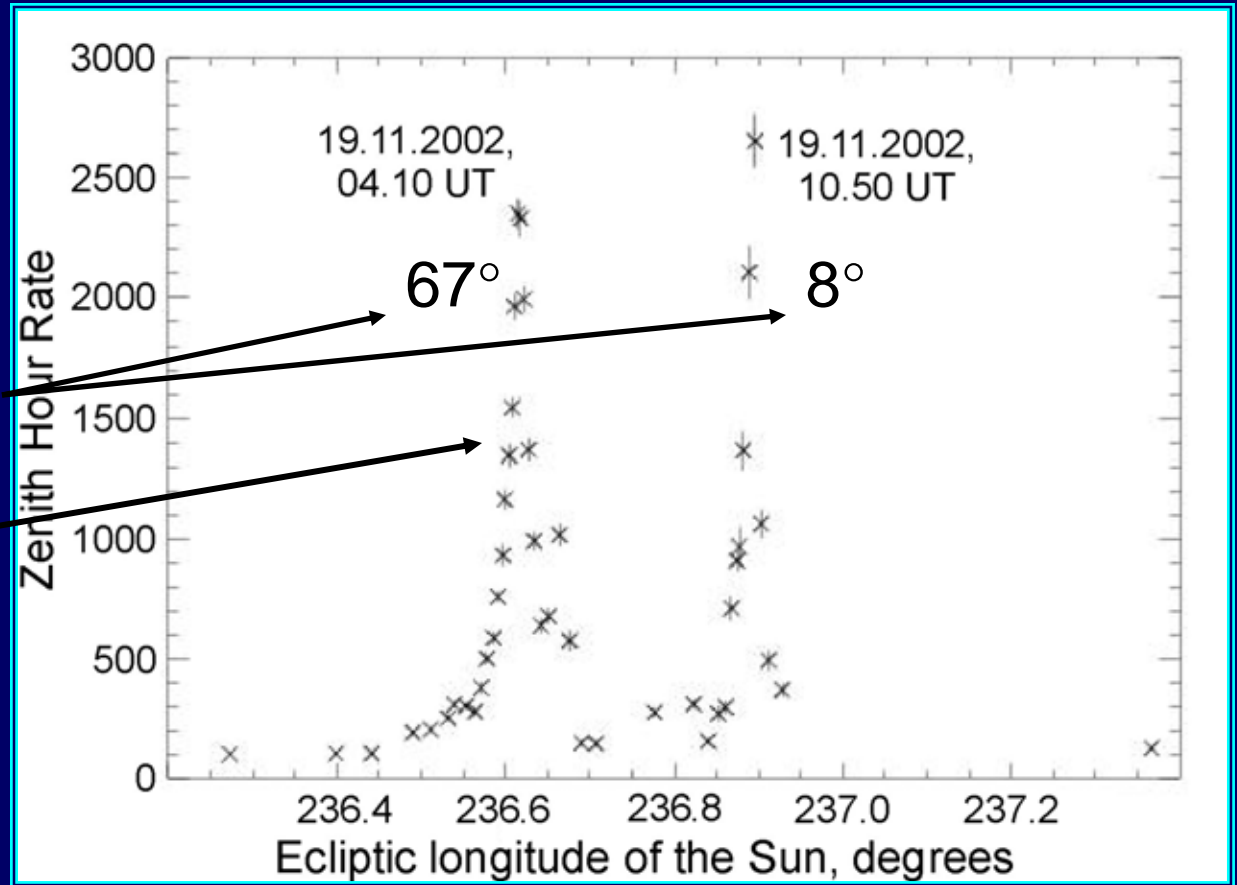
Radiant: Leo constellation

Last strong maxima:
November 1998-2002



Leonids in 2002

Last strong maxima.

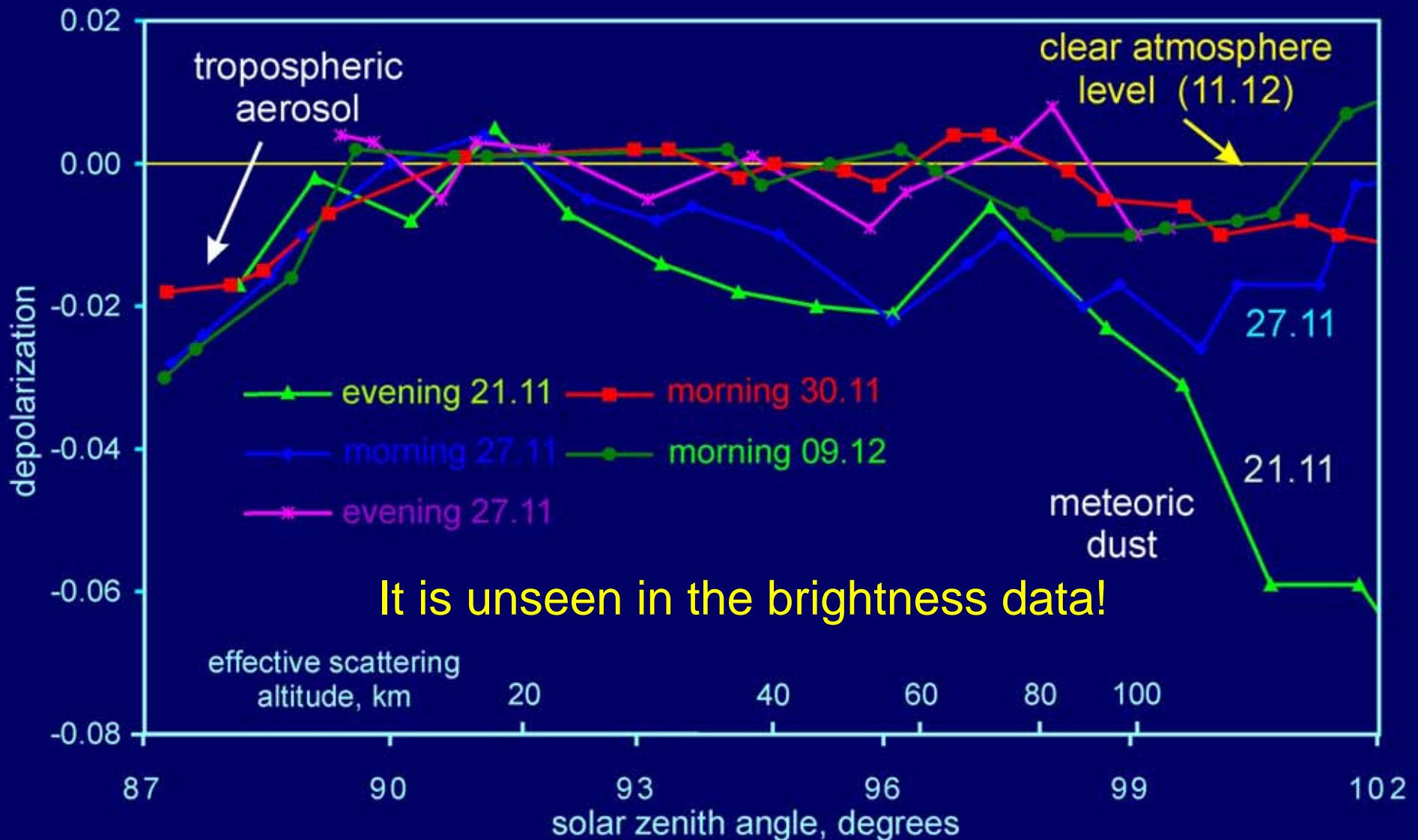


Radiant altitude in
the observation point

Meteor Dust Inflow

Data of International Meteor Organization

Twilight sky polarization – November, 2002



It is unseen in the brightness data!

Resume

Twilight sky background polarization analysis allows:

1. *To separate the multiple, molecular and aerosol scattering;*
2. *By the measurements during the whole twilight period – to find the altitude dependency of scattering matrix including the stratosphere and mesosphere;*
3. *By the wide-angle measurements – to find the angular dependency of the matrix (scattering function);*
4. *By the spectral measurements – to make the exact measurements of trace gases (such as NO_2) at different altitudes;*
5. *And making this all – to know A LOT about the atmosphere!*

Perspectives

*Multi-point (network)
twilight background measurements:*



*Dynamical investigations of aerosol
from troposphere to mesosphere
above the large territories.*

Thank you for the attention!